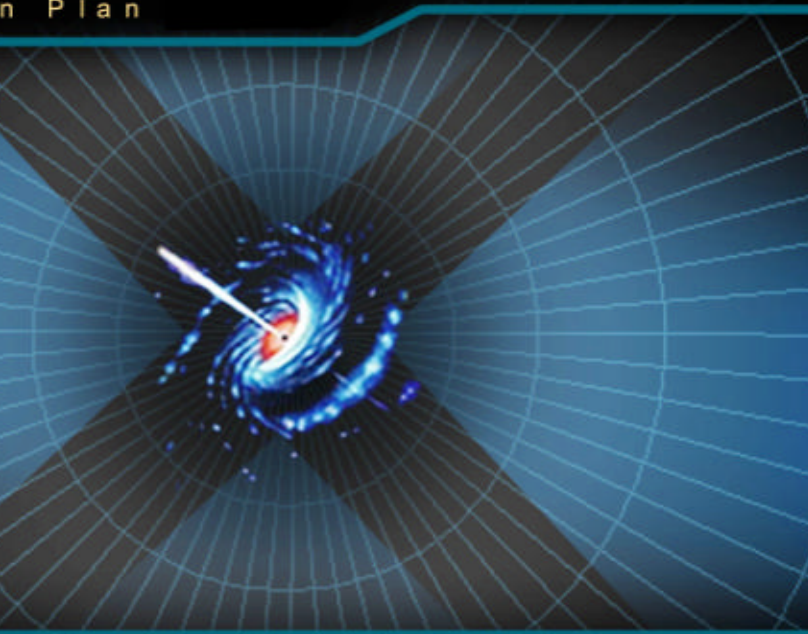


# Constellation

The Constellation X-ray Mission



## ►► TRIP Highlights

*Jean Grady*  
*Constellation-X Deputy Project Manager*  
*Goddard Space Flight Center*

## Technology Readiness and Implementation Plan (TRIP) Overview

- **Schedule of Events**

- Received Call for Reports from HQ early October 2002
- Report provided to HQ on February 3, 2003
- Received follow-up questions on Feb. 21 and March 5, 2003
- Held site visit at GSFC on March 20, 2003
- TRIP Review Report issued April 22, 2003

- **Report Content**

- Science and Instrumentation description
- Mission Implementation
- Technology Development Plans
- Management and Costs

- **TRIP Report available on the Constellation-X Project Library Web site**

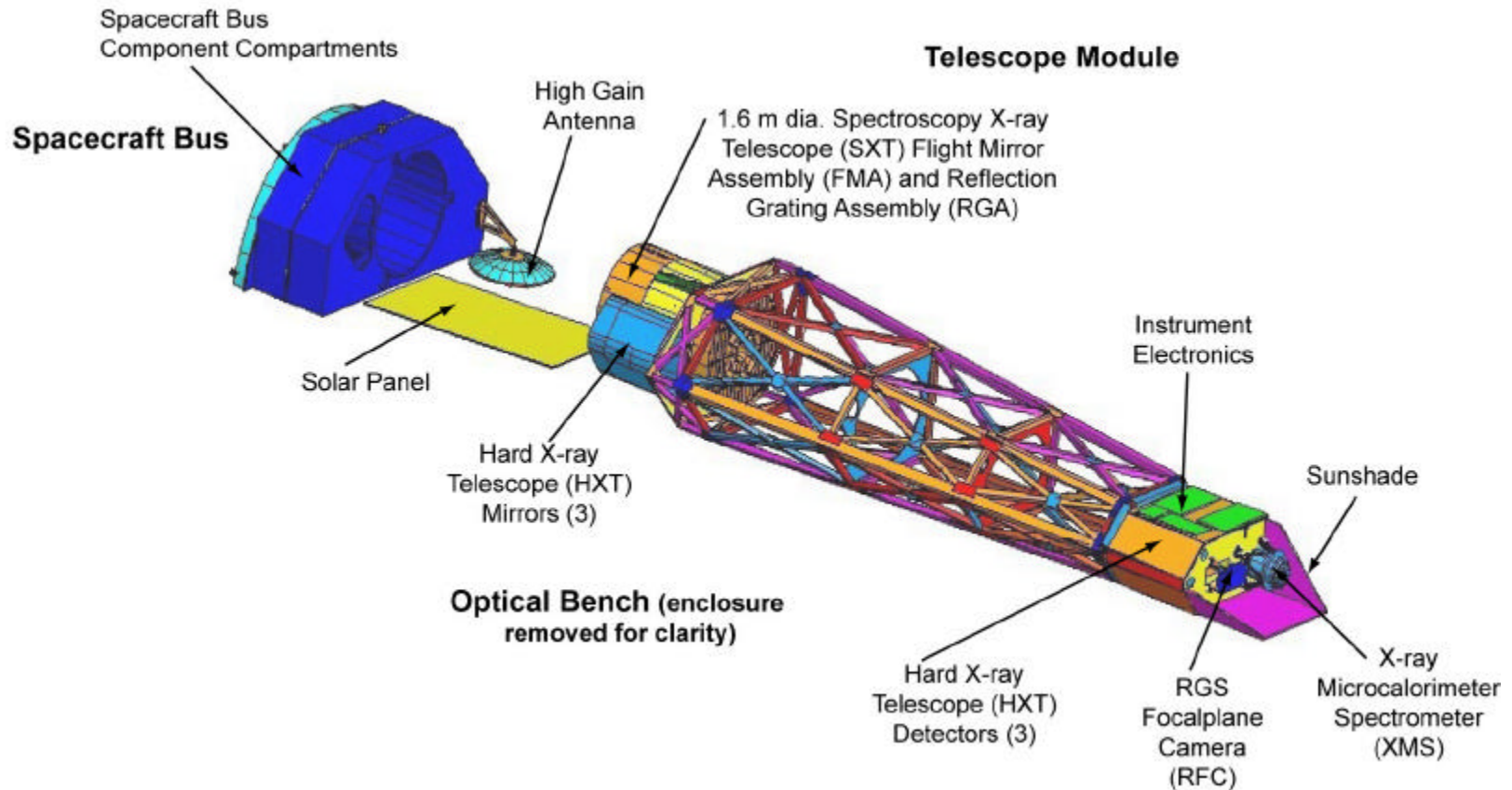
- <http://conxproject.gsfc.nasa.gov/engn.htm>

***Many thanks to all who contributed to the  
Constellation-X TRIP Report and to the TRIP  
site visit!***

## Reference Mission Configuration

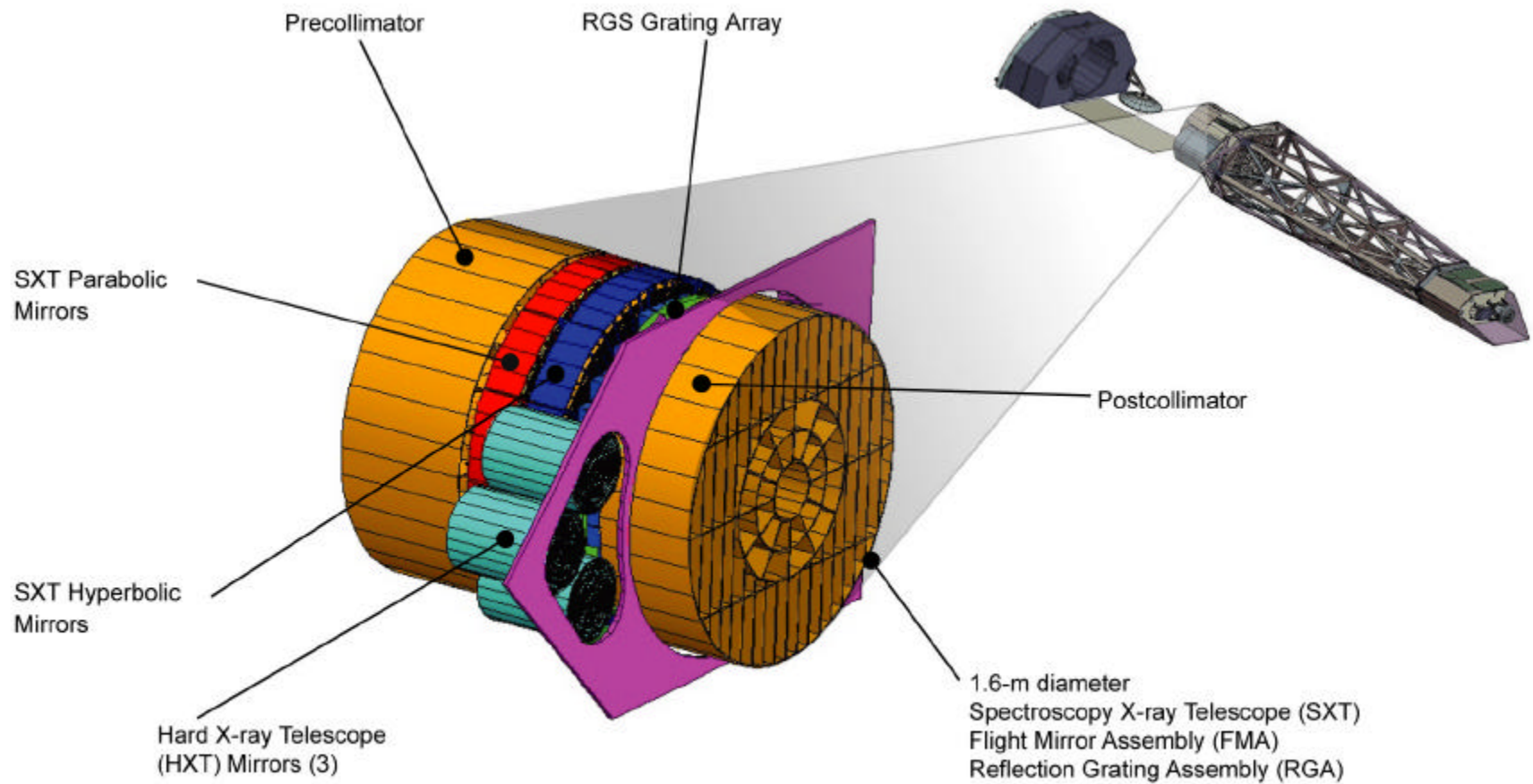
- Overall Reference Mission definition remains unchanged
  - 4 identical observatories, 2 observatories per launch, L2 orbit, etc.
- Refined modular concept for Reference Mission Configuration
  - Telescope Module, consisting of
    - Optics Module (OM)
    - Optical Bench (OB)
    - Focal Plane Module (FPM)
  - Spacecraft Bus
- Refined instrument and spacecraft block diagrams and conceptual designs
- Updated power and mass budgets
- Developed I&T flow, which takes advantage of modular design to minimize schedule and risk
- Refined Missions Operations Concept

## Exploded View of Constellation-X Observatory

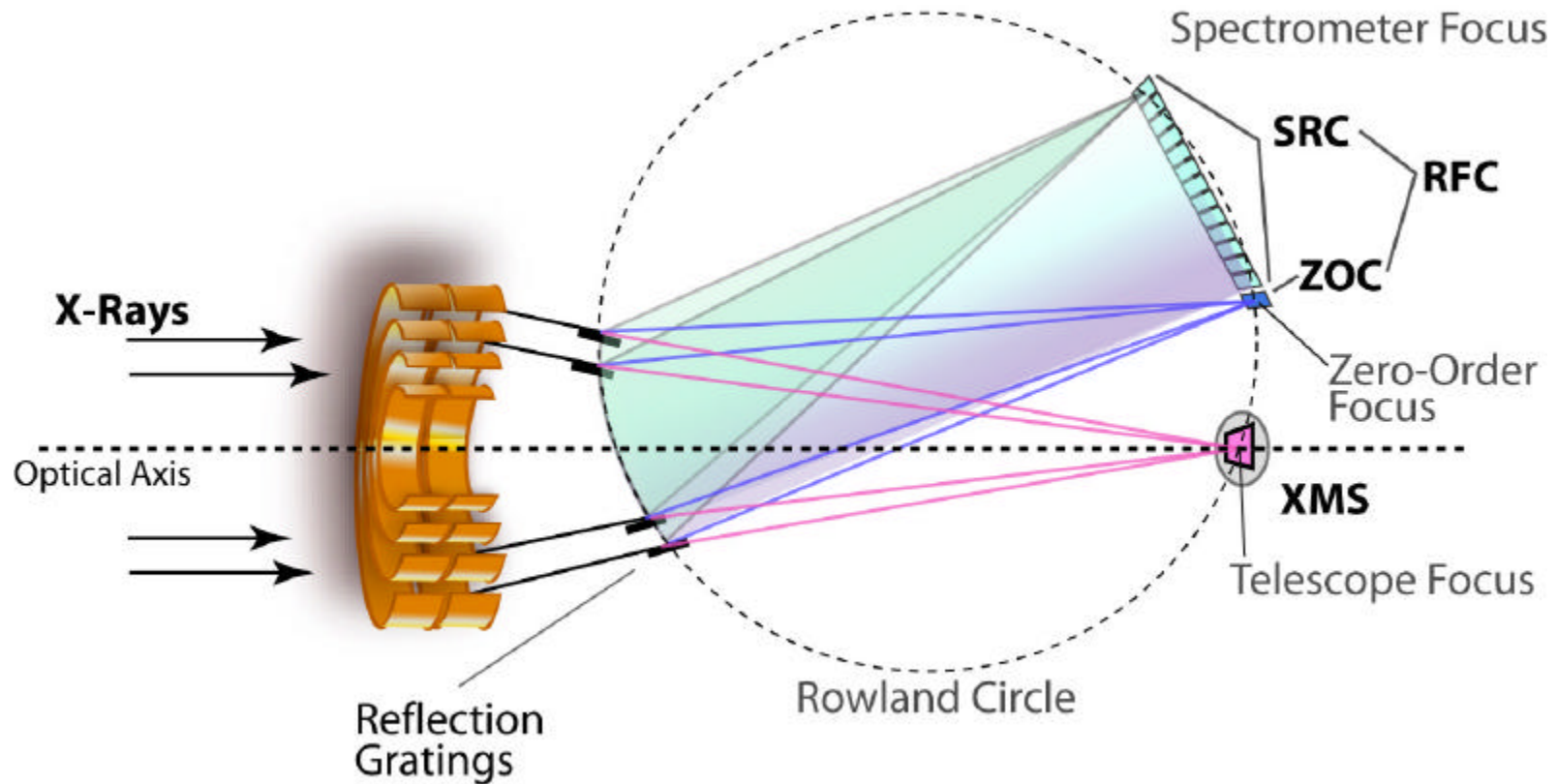




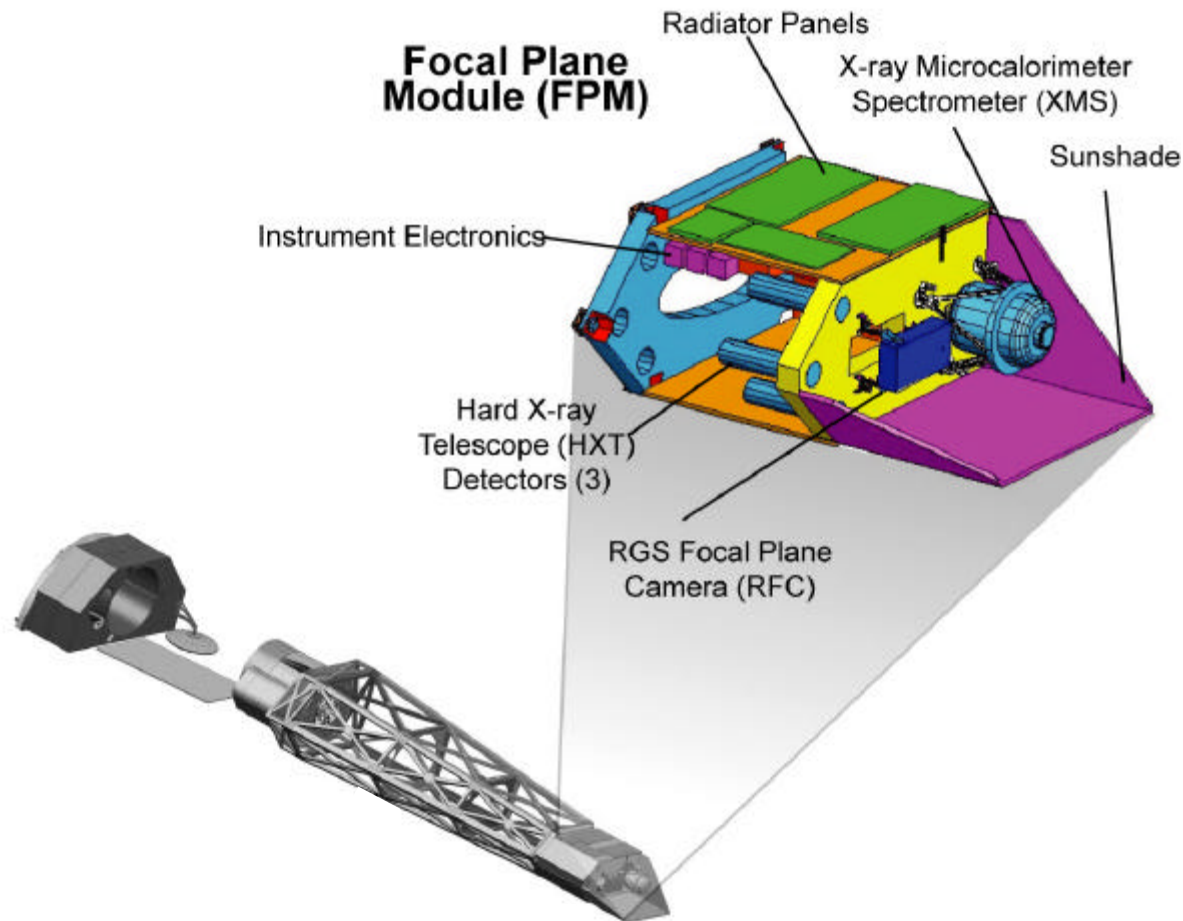
## Constellation-X Observatory — Optics Module



## SXT Optical Path



## Constellation-X Observatory — Focal Plane Module





## Telescope Module (TM) – Organization

### 2 Protective Covers/Mechanisms

Outer Protective cover (on bus)  
Inner Protective cover (not shown)  
Redundant drives – used one time

### Optical Bench

Establishes gross focal length  
Maintains alignment  
Supports baffles and cabling

### 2 Focus Adjust Mechanisms

Relaxed requirements (1 mm)  
Redundant – used infrequently

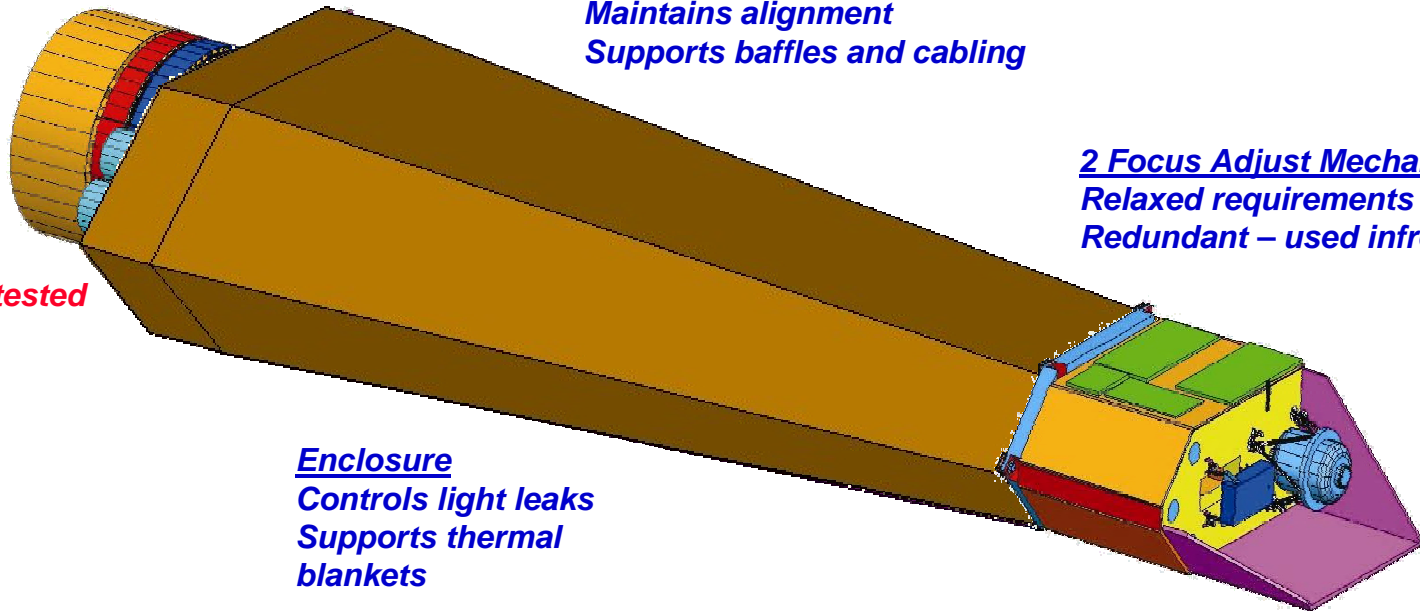
**Optics Module**  
(integrated and tested separately)

### Enclosure

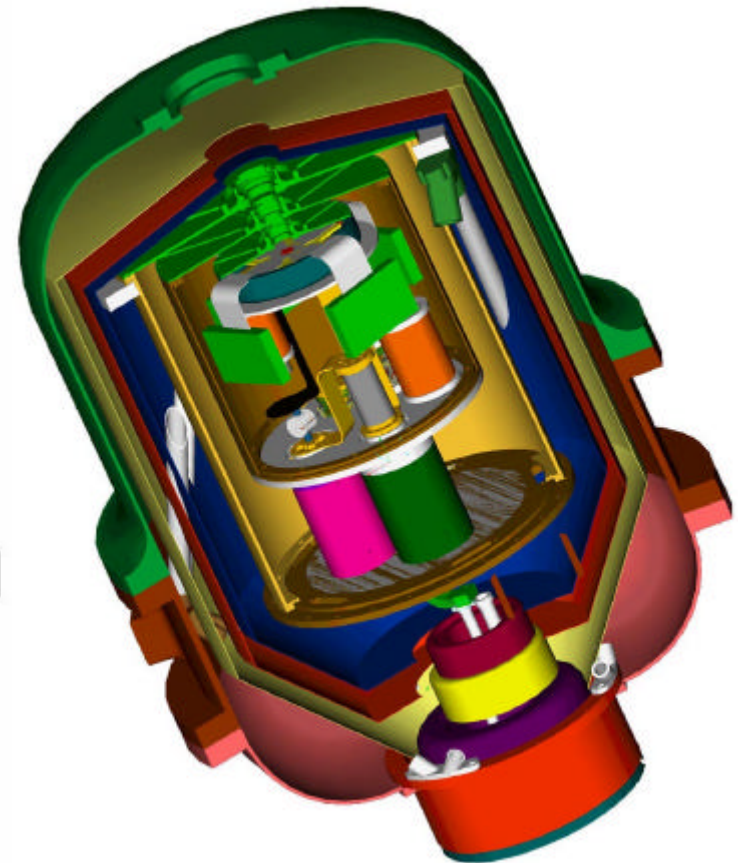
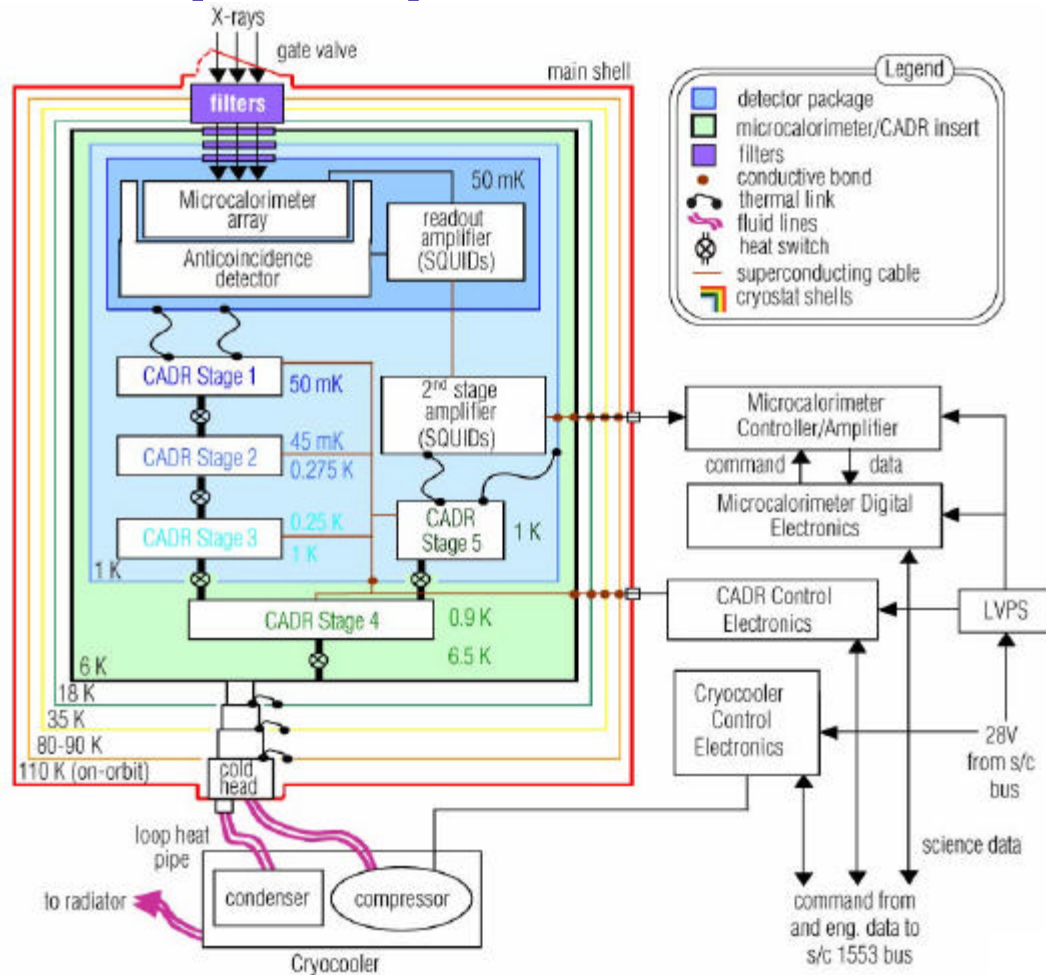
Controls light leaks  
Supports thermal blankets

**Calibration Source**  
**Blocker Mechanism**

**Focal Plane Module**  
(integrated and tested separately)



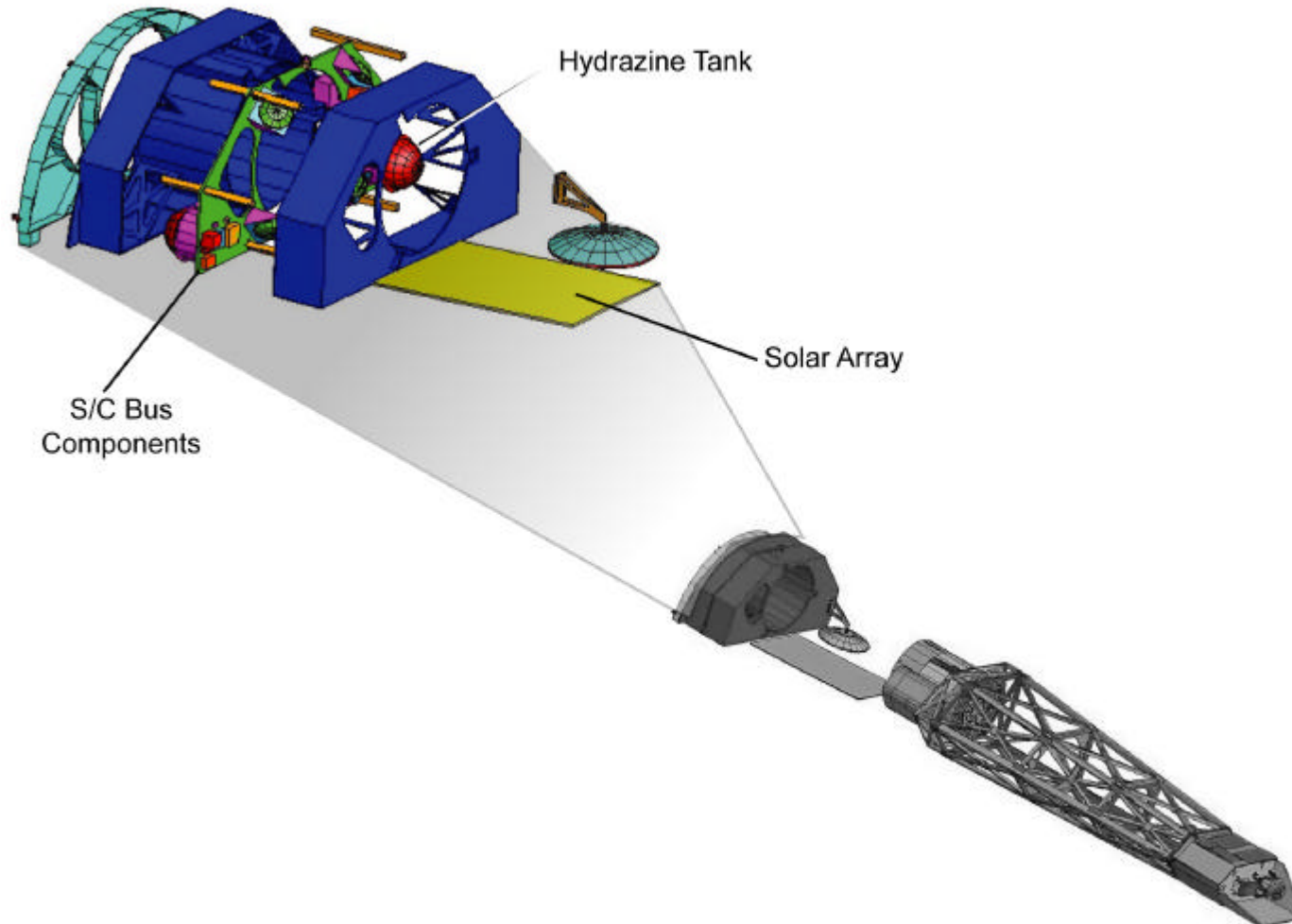
# X-ray Microcalorimeter Spectrometer (XMS) Block Diagram and Conceptual Implementation



Size ~ 50 x 75 cm

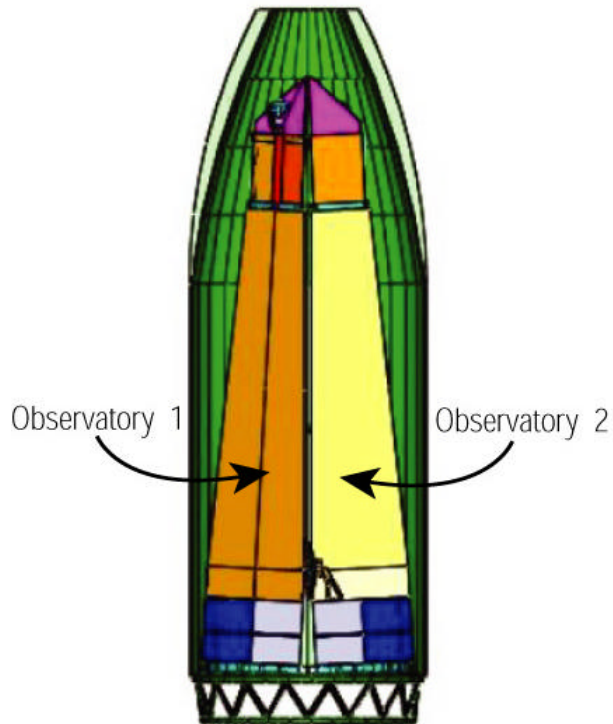
Mass ~ 150 kg, including electronics

## Constellation-X Observatory — Spacecraft



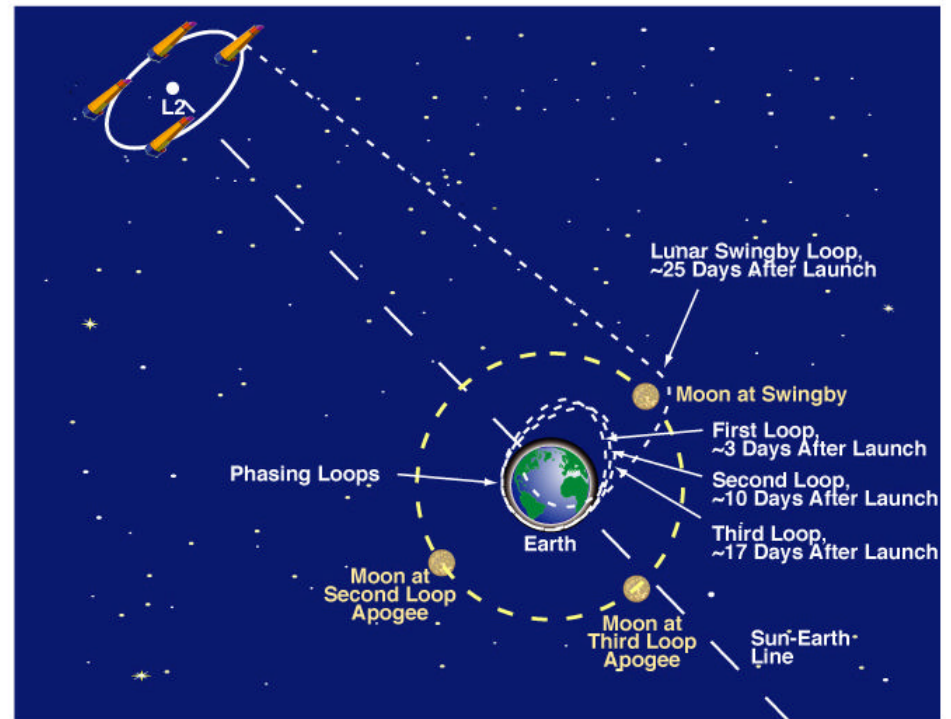
# Constellation-X Launch Configuration and Orbit

## Constellation-X Launch Configuration



ATLAS V

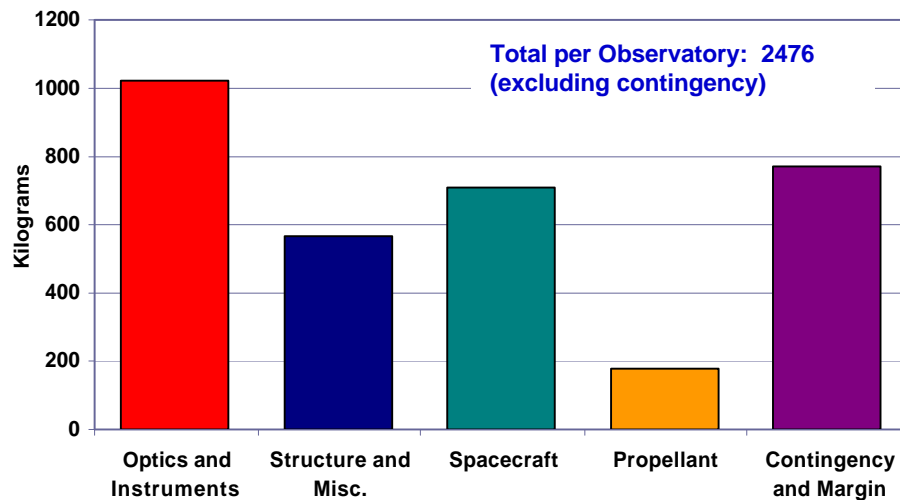
## Trajectory with Phasing Loops and Lunar Swingby





## Mass and Power Resources

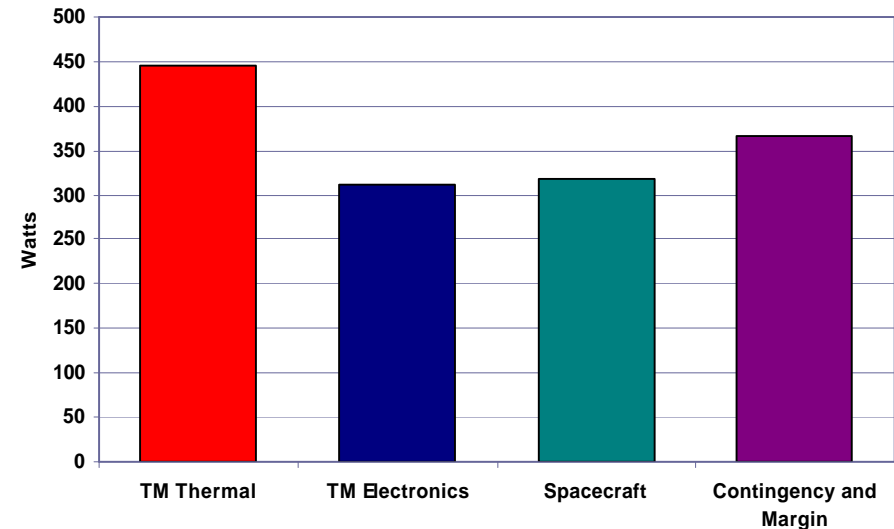
**LV Performance = 6498 Kg**



**% Contingency & Margin = 34%**

**Mass**

**Solar Array EOL = 1442 W**



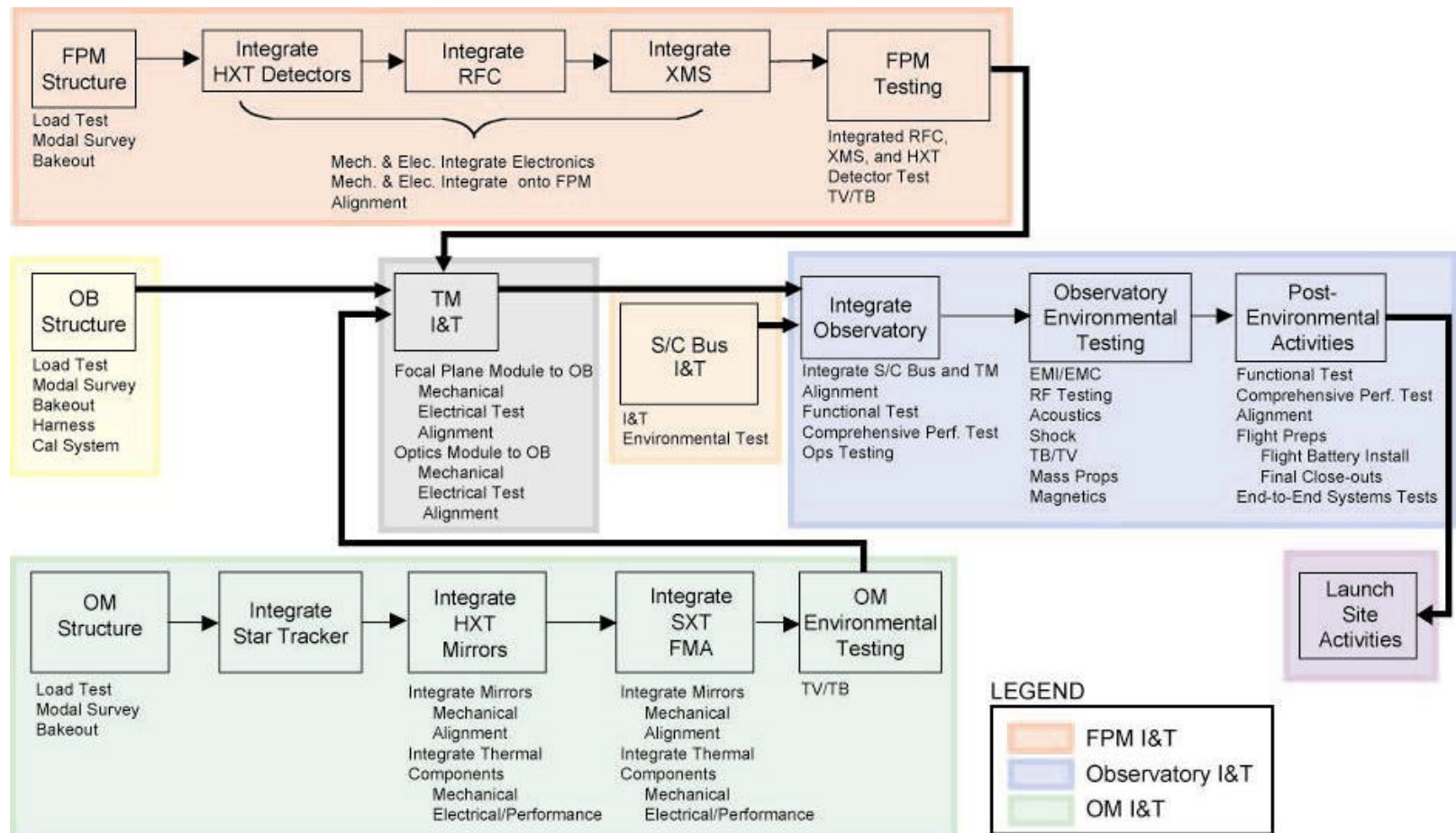
**% Contingency & Margin = 34%**

**Power**

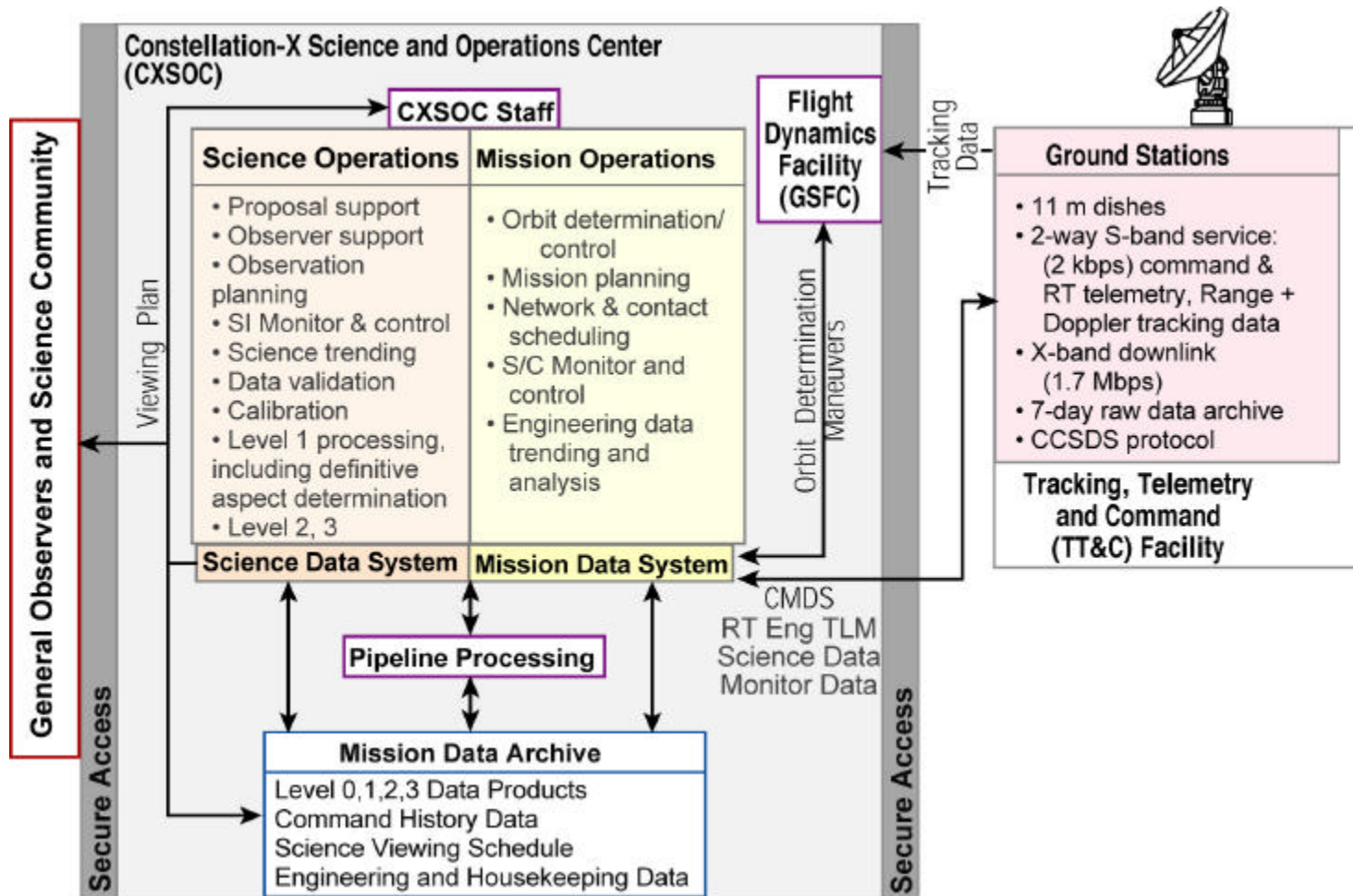
- Sufficient resources are available for the mission



## I&T Flow



# Constellation-X Mission Operations



## Requirements and Error Budgets

- Mission Top Level Requirements remain stable
  - Incorporated Spectral Resolving Power of 1500 from 6 to 10 keV
- Updated Mission and Instrument error budgets
- Documented flow down of requirements to each instrument and instrument elements

## Key Top Level Requirements

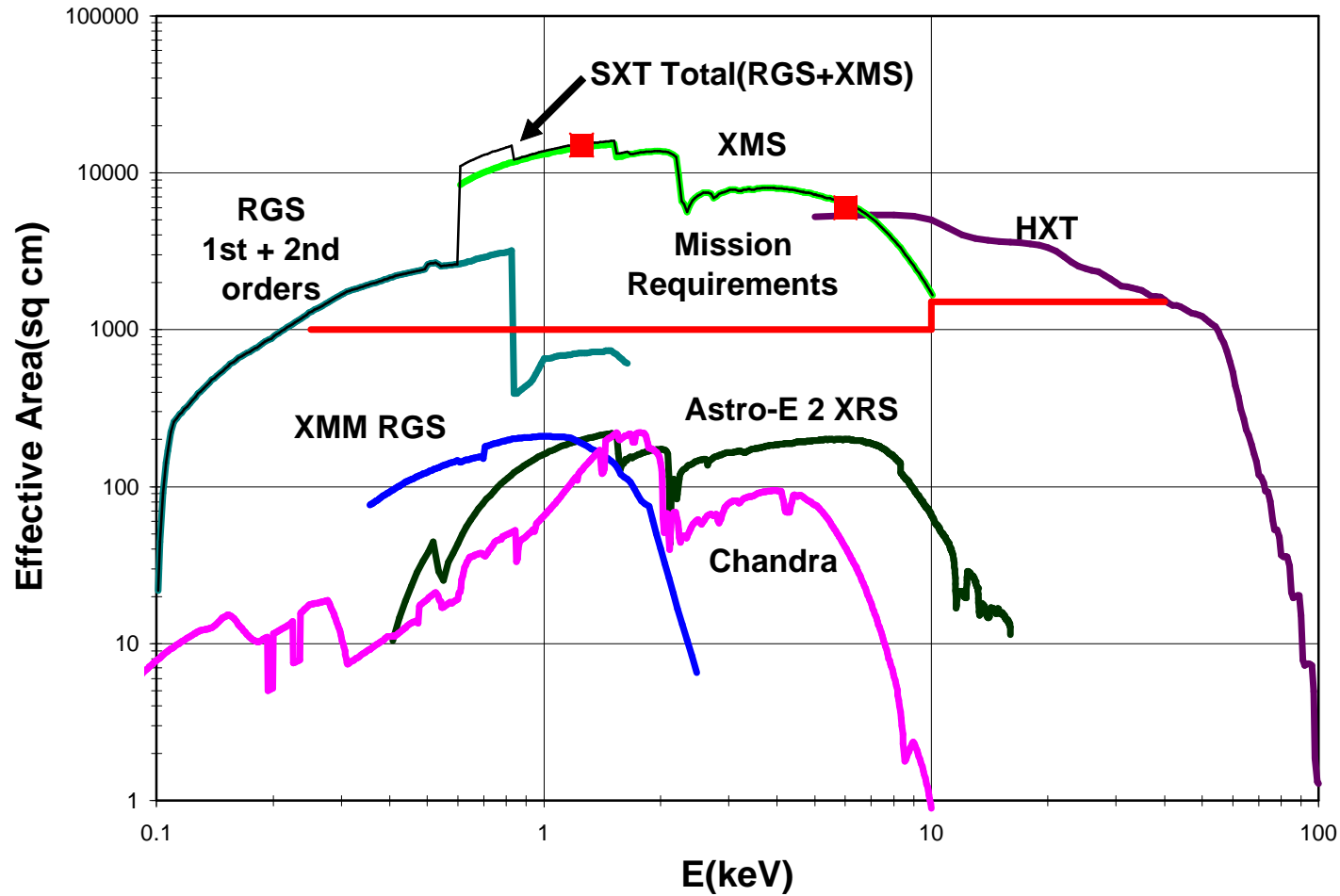
Parameter	Baseline Requirement
<b>Bandpass</b>	<b>0.25 to 40 keV</b>
<b>Effective Area</b>	
0.25 to 10 keV	1,000 cm <sup>2</sup>
1.25 keV	15,000 cm <sup>2</sup>
6.0 keV	6,000 cm <sup>2</sup>
10 to 40 keV	1,500 cm <sup>2</sup>
<b>Spectral Resolving Power (E/DE)</b>	
0.25 to 6.0 keV	300
6 to 10 keV	1,500
10 to 40 keV	10
<b>Angular Resolution (HPD)</b>	
<10 keV	15 arcsec
>10 keV	1 arcmin
<b>Fields of View</b>	
<10 keV	2.5 arcmin
>10 keV	8 arcmin
<b>Bright Source Limit</b>	<b>40,000 cps</b>
<b>Absolute Timing (relative to UTC)</b>	<b>100 microsec</b>
<b>Mission Lifetime</b>	<b>4 years at full capacity</b>

# SXT Effective Area Budget

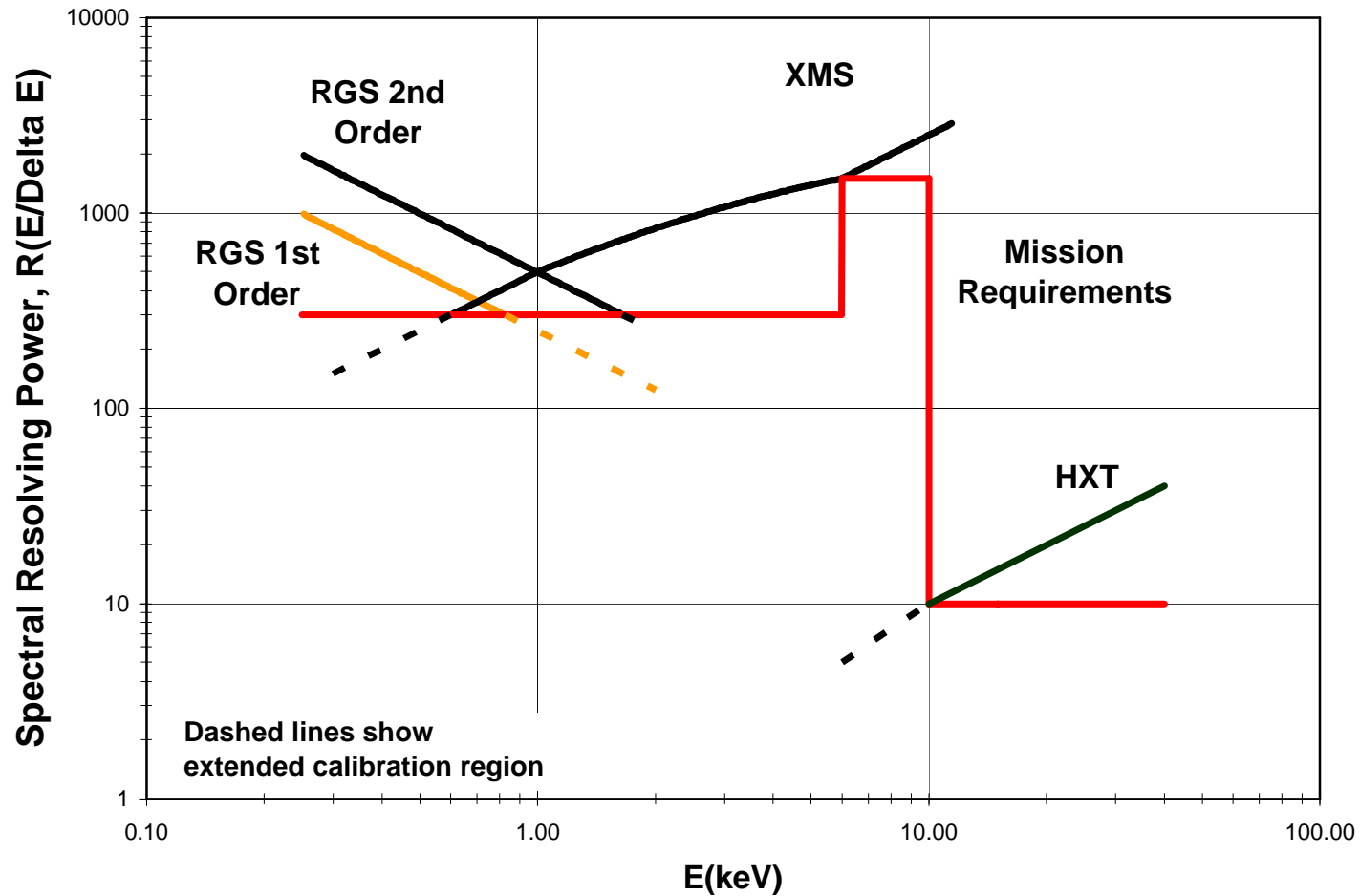
SXT Effective Area Budget				
	Area At Energy			Comment
	0.25 keV	1.25 keV	6keV	
SXT FMA Geometric Area	59,400	59,400	59,400	230 shell design
Reflectivity loss	-17,118	-18,641	-50,324	Gold coating
SXT FMA Effective Area	42,282	40,759	9,076	9076 was 8709 in TRIP (error)
SXT Effective Area Losses				
Structural blockage	-5,919	-5,747	-1,534	Baseline housing design
P-H Shell Alignment	-423	-611	-182	1% loss
Aperture Alignment	-211	-306	-91	1/2% to 1% loss (energy dependent)
SXT Contamination - EOL	-423	-408	-91	1% (Chandra experience)
SXT Effective Area	35,306	33,687	7,179	
(per SXT)	8,826	8,422	1,795	
XMS Area Feed	19,627	18,826	6,109	
RGS Area Feed	15,678	14,861	1,070	
Instrument/Telescope losses				
RGS Internal Vignetting	-784	-743	-54	
XMS(Cal QE, Filter, fill factor)	-19,627	-3,212	-410	
RGS(Grat Effy, CCD QE, Filter)	-12,659	-13,280	0	
Grating internal alignment	-157	-149	-11	1%
Off-axis operation	-14	-172	-68	Raytrace
Inst Contamination - EOL	-784	-941	-14	0.2% to 5% energy dependent
Total Area - Predicted	1,281	15,191	6,622	
Total Area - Requirement	1,000	15,000	6,000	
Margin(%)	28.1	1.3	10.4	



## Mission Effective Area



## Mission Spectral Resolving Power



# SXT Angular Resolution Error Budget

Item (HPD – arcsec)	Rqmt	Margin	Allocation/Predictions				Rationale
RGS Resolution	15.00	4.01	14.46				4 satellites, post-processed
Co-add 4 satellites				1.00			Superposition of data using X-ray centroids
On-Orbit Telescope - single satellite			14.42				RSS
CCD pixelization error				0.41			0.5 arcsec pixels
▪ Grating resolution errors				5.00			Estimate
XMS Resolution	15.00	4.95	14.16				4 satellites, post-processed
Co-add 4 satellites				1.00			Superposition of data using X-ray centroids
On-Orbit Telescope - single satellite			14.12				RSS
▪ Calorimeter pixelization error				4.08			5 arcsec pixels
▪ Telescope level effects				5.20			RSS
– Image reconstruction errors (over obs)					4.24		RSS
– SXT/Telescope mounting strain					2.00		Eng. estimate based on Chandra experience
– SXT/SI vibration effects					2.00		Chandra experience (jitter)
– SXT/SI misalignment (off-axis error)					1.00		Chandra experience
– SXT/SI focus error					0.20		Analysis
▪ SXT Optics - on-orbit performance				12.48			RSS
– SXT Mirror launch shifts					2.00		Eng. est. based on Chandra
– Thermal errors					2.24		RSS
– Material stability effects					1.00		Est. based on Chandra work
– SXT Mirror, as built					12.07		RSS
➤ Gravity release						1.50	FEA analysis using vertical assy
➤ Bonding strain						3.00	Eng. estimate, analysis in process
➤ Alignment errors (using CDA)						3.38	RSS
➤ Installation in housing						5.00	Est. based on OAP1 testing
➤ Optical elements						9.90	Est. based on tech dev program

Legend:

Requirement

Margin

RSS Prediction

Allocation

# SXT FMA Requirements

SXT FMA Performance Requirements		Trace to Top-Level Mission Requirements
Bandpass	0.25 to 10 keV	Allocation of mission bandpass to SXT
Effective area (per mirror) @0.25 keV @1.25 keV @6 keV	8,826 cm <sup>2</sup> 8,421 cm <sup>2</sup> 1,722 cm <sup>2</sup>	Provides 33,000 cm <sup>2</sup> at 1 keV and 6,900 cm <sup>2</sup> at 6 keV for the mission. Allows effective area losses due to detector efficiency, etc., to achieve TLRD baseline requirement per error budget summarized in Table 1-2.
Angular resolution	12.5 arcsec HPD	Error budget allocation to mirror that allows telescope system to achieve requirement of 15 arcsec with 4 arcsec margin combined by RSS (Table 1-3).
Field of view	2.5 arcmin	Exceeds instrument FOV; defined by detector FOV
Derived Requirements: SXT Mirror		Derivation
Diameter	1.6 m	To meet mission area requirements with 4 mirrors
Focal length	10 m.	Consistent with grazing angle requirements for 1.6 m diameter mirror.
Axial length	<70 cm	To fit within envelope and meet fabrication considerations
Operating temperature	20±1° C nominal	Range is per allocation from SXT angular resolution error budget (Table 1-3); minimizes angular distortions imposed by temperature change to components. Operating temperature is determined by optics assembly temperature
Mass	642 kg	Current engineering estimate
Derived Requirements: SXT Grating: See Table 1-3		
Derived Requirements: Thermal Pre/Post collimators		
Temperature gradient	1° C across diameter 1° C axial	Allocation from SXT angular resolution error budget (Table 1-3); minimizes angular distortions imposed by temperature gradients
Mass	47 kg	Current engineering estimate

## RGS System Level Requirements

RGS Performance Requirements		Trace to Mission Top-Level Requirements
Bandpass	0.25-2.0 keV (6 to 50 Å)	In combination with XMS, meets spectral resolution reqmts over the 0.25 – 10 keV bandpass. 1 to 2 keV used for calibration with XMS
Spectral resolving power, R (???)	$\geq 300$ below 1 keV	Meets TLRD baseline requirement for R
Effective Area @0.25 keV @0.6 keV @1.25 keV	250 cm <sup>2</sup> 625 cm <sup>2</sup> 175 cm <sup>2</sup>	Flowdown from mission baseline effective area requirement



# RGA Requirements

Derived RGS Grating Array Requirements		Derivation
Grating efficiency: @0.25 keV (1st Order) @0.6 keV (1st Order) @1.25 keV (2nd Order)	>0.14 >0.22 >0.06	Flowdown from area requirements. Theoretical efficiency with 50% margin. Met with 40% margin when measured efficiencies for anisotropically etched grating test ruling are used
Interception factor	0.57	Fraction of X-rays entering RGA intercepted by gratings and dispersed in the various orders. Flowdown from area requirements
Straight-through factor	0.38	See Inteception factor (above)
Grating goove parameters a: incidence angle ?: graze angle d: groove spacing	a = 1.61 deg. ? = 2.21 deg. 1/d = 407 mm <sup>-1</sup>	Given 15" HPD telescope, and requiring = ??? 400 at blaze (blaze = 0.605 deg. reflectivity is optimized there using scalar diffraction theory.
Grating flatness	≤2 arcsec FWHM	Grating error budget flowdown for spectral resolution. Combined with alignment error, allows broadening of the line spread function core by no more than 30% and SXT mirror dominates
Grating to grating alignment	≤2 arcsec FWHM	See grating flatness item (above)
Mass	50 kg	Current engineering estimate

# RFC Requirements

Derived RGS Focal Plane Camera Requirements		Derivation
Quantum Efficiency @0.25 keV @0.6 keV @1.25 keV	>0.86 >0.93 >0.98	Flowdown from area requirements
Energy Resolution at 250 eV	> 90% events within 100 eV band	Required to separate spectra from overlapping orders. The requirement is met with 20% margin by state-of-the-art (ACIS-S) BI CCD's
Optical Blocking Filter -Visible light rejection	>10 <sup>6</sup>	Optical light rejection to avoid CCD pulse height confusion
X-ray transmission @0.25 keV @1.25 keV	>0.8 >0.98	Flowdown from area requirements in conjunction with grating efficiency meets the top-level area requirements
Optical starlight rejection	≤1 electron/pixel/readout for 10 magnitude star	Joint requirement on pre-collimator, SXT straylight performance, and SRC CCD optical blocking filter performance
Pixel size	24 microns	Required to critically sample the Point Response Function
SRC number of pixels, dispersion direction	1.3 X 10 <sup>4</sup>	Required to cover the dispersed instrument bandpass (0.25 to 2 keV), given above pixel size and SXT focal length. (1024 pixels x 13 CCDs)
SRC number of pixels, cross-dispersion direction	512	Required to provide adequate areas to enable background subtraction
ZOC CCD format	1024 X 1024	Identical to SRC chips to minimize costs
Frame readout rate	2 second integration time per frame	< 50% pileup in central CCD pixel for bright source limit, assuming 20% flux in single emission line
Operating temperature	-60° C to -80° C	Reduces hot and flickering pixels
Mass	33 kg	Current engineering estimate

## Technology Development

- Updated Technology Roadmaps to achieve TRL 6
- Defined Technology Gates:
  - Subsets of TRL demonstrations defined in the TRIP report and
  - Represent significant improvements in performance or scale
- Summarized Technology Development milestones, including goals of each technology demonstration stage for each technology

# Technology Development Roadmap Summary


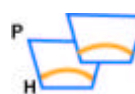

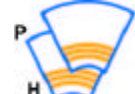


System	Technology	Heritage	Required Improvement	Req't	Subsystem Technology Readiness Level by Fiscal Year				
					1998	Current	2004	2005	2006
FMA	SXT Mirror	Astro-E/E2, BBXRT, ASCA	Angular resolution	12.5 arcsec	TRL 2	TRL 3-4	TRL 4	TRL5	TRL 6
		XMM-Newton	Larger diameter	1.6 m					
RGS	Gratings (RGA)	XMM-Newton, Chandra	Low mass	0.2g/cm <sup>2</sup>	TRL 3	TRL 3	TRL 5		TRL 6
			Mass production	25/day					
	CCD Detector (RFC)	Chandra, ASCA	Production yield	20%	TRL 2	TRL3	TRL 4	TRL 6	
			Event drive						
XMS	Microcalorimeter	Astro-E/E2	Larger array	32 x 32 pixels	TRL 3	TRL 4	TRL 5	TRL 6	
			Energy resolution	4 eV					
	ADR	Astro-E/E2 HAWC, XQC	Warmer sink	6 K	TRL 3	TRL 4		TRL 5	TRL 6
			Cont. operations						
	Cryocooler	HST, TES, AIRS	Lower temperature	6 K	TRL 3	TRL 4		TRL 5	TRL 6
HXT	HXT Mirrors	HEFT, InFOC $\mu$ S	Angular resolution	60 arcsec	TRL 3	TRL 4	TRL 5	TRL 6	
	HXT Detectors	HEFT, Swift	Low energy response	6 keV	TRL 3	TRL 4-6	TRL 5	TRL 6	

# Technology Gates

System	Technology	Performance Parameter	State-of-the-Art	Current	Technology Gates			Nominal Flight Requirement
					2004	2005	2006	
SXT FMA	Mirrors	Angular Resolution (HPD)	80 arcsec	<15 arcsec (reflector only)	12.5 arcsec	12.5 arcsec		12.5 arcsec
		Diameter	40 cm	20 cm	50 cm	160 cm		160 cm
RGS	Gratings (RGA)	Mass/unit area	0.6g/cm <sup>2</sup>	—	0.2g/cm <sup>2</sup>		0.2g/cm <sup>2</sup>	0.2g/cm <sup>2</sup>
		Groove density/variation	645 lines/mm / 7%	500 lines/mm / 0%	N/A		407 lines/mm/5%	407 lines/mm/5%
	CCD	Quantum Efficiency	15%	25%		>80%		86%
XMS	Microcalorimeter	Array Size	32 pixels	25 pixel array 4 pixel readout	64 pixel array 16 pixel readout	1032 pixel array 96 pixel readout		1032 pixel array 1024 pixel readout
		Energy Resolution	4.8 eV at 6 keV	10 eV at 6 keV		4 eV at 6 keV		4 eV at 6 keV
	ADR	Cold/Hot end	50mK/1.1K	50mK/4K		50mK/6K		50mK/6K
		Operating mode	Periodic	Continuous		Continuous		Continuous
	Cryocooler	Cooling power	1.5 mW at 55K	0 mW at 5.4K		20 mW at 6K		20 mW at 6K
HXT	Mirrors	Obscuration throughput	40%	60%		75%		75%
	Detectors	Low Energy Threshold	~17 keV	~17 keV		6 keV		6 keV



# SXT FMA Technology Development Roadmap Summary

	Optical Pathfinder Assembly		Engineering Unit	Mass Alignment Pathfinder	Prototype	
	OAP #1	OAP #2				
Configuration						
Module Type	Inner	Inner	Inner	Inner	Outer	Wedge (2 Outer & 1 Inner)
Housing Material	Aluminum	Titanium	Composite	Composite	Composite	Composite
Focal Length	8.4 m	8.4 m	8.4 m	8.4 m	10.0 m	10.0 m
Reflector Length (P&H)	2 x 20 cm	2 x 20 cm	2 x 20 cm	2 x 20 cm	2 x 20-30 cm	2 x 20-30 cm
Nominal Reflector Diameter(s)	50 cm	50 cm	50 cm±	50 cm±	160 cm 120 cm± 100 cm	160 cm±40 cm± 120 cm± 100 cm±
Goals	<ul style="list-style-type: none"> <li>Align 1 reflector pair (P&amp;H)</li> <li>Evaluate mirror assembly design, alignment and metrology</li> </ul>	<ul style="list-style-type: none"> <li>Align 1 reflector pair</li> <li>Evaluate reflector</li> <li>Evaluate mirror bonding</li> </ul>	<ul style="list-style-type: none"> <li>Align up to 3 reflector pairs to achieve &lt;12.5 arcsec</li> <li>Eval. assembly gravity sag</li> <li>X-ray and environmental test</li> <li>Evaluate composite housing</li> </ul>	<ul style="list-style-type: none"> <li>Align 3 reflector pairs</li> <li>Evaluate tooling and alignment techniques for mass production</li> <li>X-ray test</li> </ul>	<ul style="list-style-type: none"> <li>Flight-like configuration outer module</li> <li>Environmental and X-ray test</li> <li>Largest reflectors</li> </ul>	<ul style="list-style-type: none"> <li>Demonstrate largest and smallest diameter reflectors</li> <li>Demonstrate module to module alignment</li> <li>Environmental and X-ray test</li> </ul>
TRL	TRL 3		TRL 4		TRL 5	TRL 6
Timeframe	Q2 of FY03	Q3 of FY03	Q1 of FY04	Q1 of FY05	Q4 of FY05	Q4 of FY06
Technology Gate			◆		◆	

# RGA Technology Development Roadmap Summary

Parameters	State-of-the-Art XMM-Newton	Grating Fab Demo	Large Area Grating	Demo Assembly	Grating Chirp	Grating Module Engineer Unit	Grating Module Flight Unit
Substrate Flatness	<2 arc sec	~30 arc sec	<2 arc sec	<2 arc sec	<2 arc sec	<2 arc sec	<2 arc sec
Grating Size	200 x 100 mm	20 x 20 mm	140 x 100 mm (70% flight size)	200 x 100 mm (nom. flight size)	(200 x 100 mm)	(200 x 100 mm)	(200 x 100 mm)
Grating Mass/Unit Area	0.6g/cm <sup>2</sup>	—	—	0.2g/cm <sup>2</sup>	—	0.2g/cm <sup>2</sup>	0.2g/cm <sup>2</sup>
Groove Form	0.7 deg blaze	0.7 deg blaze	0.6 deg blaze	NA	0.6 deg blaze	0.6 deg blaze	0.6 deg blaze
Ruling Density/Variation	646 l/mm / 7%	500 l/mm / NA	407 l/mm / 0%	NA	407 l/mm / 5%	407 l/mm / 5%	407 l/mm / 5%
Groove Fabrication Process	Epoxy multi-gen replication of mechanically ruled master grating	Interference lithography & anisotropic etch Si (111) plane facet	Scanning Beam Interference Lithography (SBIL) Si (111) plane facet	NA	Variable Period (VP) SBIL pattern & anisotropic etch Si (111) plane facet	VPSBIL pattern & anisotropic etch Si (111) plane facet	VPSBIL pattern & anisotropic etch Si (111) plane facet
Ass'y Level & Properties	Grating Array	Single grating	Single grating	Module	Single grating	Module	Module
Gratings per Module	182 per array			3 or more gratings		~10 gratings	~10 gratings
Grating-to-Grating Align't	2 arc sec			2 arcsec		2 arcsec	2 arcsec
Other Goals	NA	<ul style="list-style-type: none"> <li>X-ray test atomically smooth groove facet</li> </ul>	<ul style="list-style-type: none"> <li>X-ray efficiency test large area grating for groove quality and uniformity</li> </ul>	<ul style="list-style-type: none"> <li>Grating substrates fab'd w mass production processes applicable to flight gratings</li> </ul>		<ul style="list-style-type: none"> <li>End-to-end X-ray test of grating module with SXT mirror segment</li> </ul>	
				<ul style="list-style-type: none"> <li>Flight representative module structure</li> </ul>		<ul style="list-style-type: none"> <li>Flight like gratings and modules</li> </ul>	
				<ul style="list-style-type: none"> <li>Verify alignment before/after environmental test</li> </ul>		<ul style="list-style-type: none"> <li>Verify alignment before/after environmental test</li> </ul>	
TRL	TRL 9	TRL 3	TRL 4	TRL 5		TRL 6	
Timeframe		Current	Q2 FY04	Q4 FY04	Q3 FY05	Q2 FY06	
Technology Gate				◆		◆	

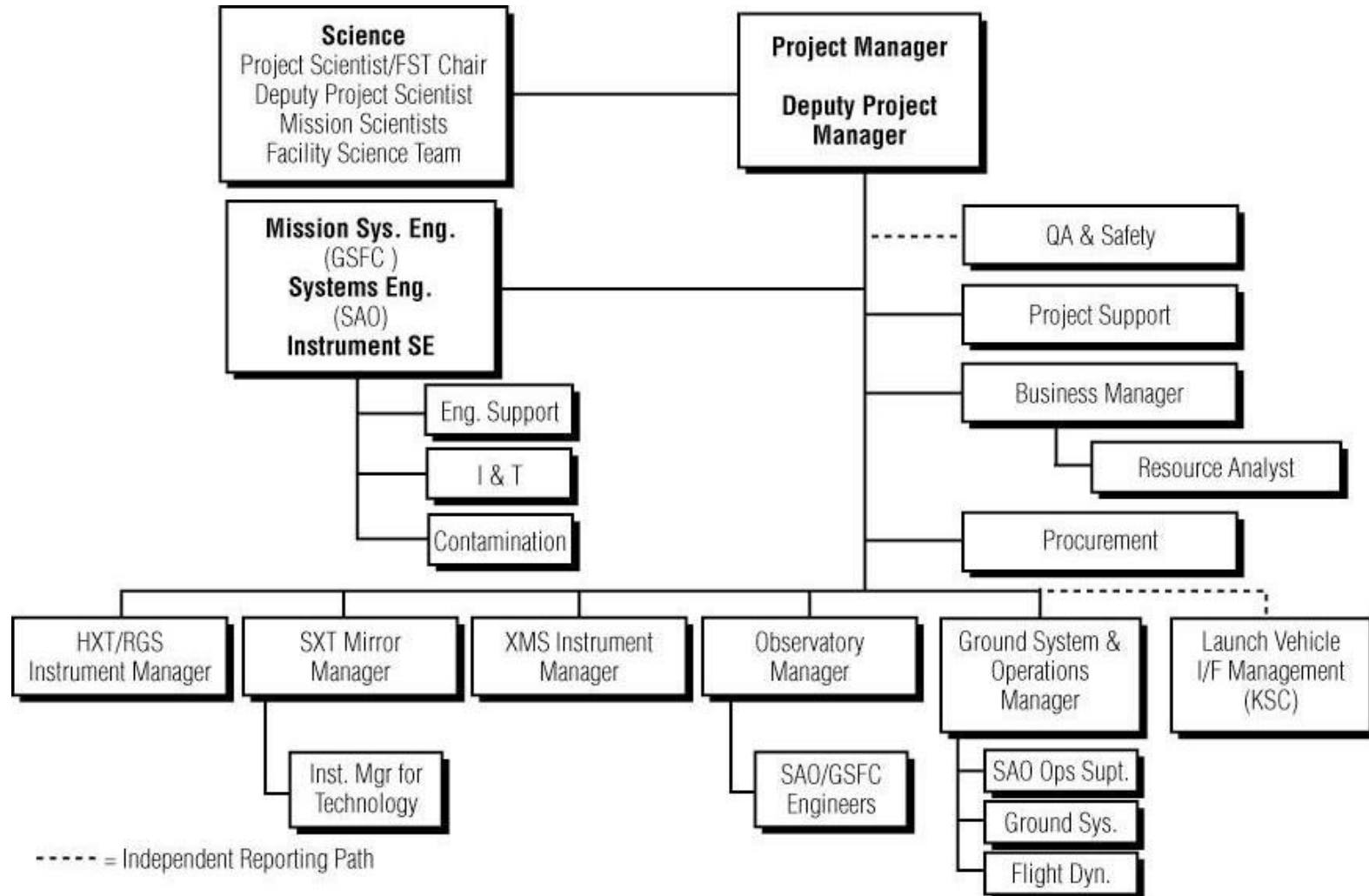
# RGS Focalplane Camera (RFC) Technology Roadmap

Parameter	State-of-the-Art Chandra ACIS	Current	ED-CCD Gen1 Lot1	ED-CCD Gen2 Lot1	ED-CCD Gen2 Lot2	ED-CCD Gen2 Lot3	Engineering Unit Focal Plane	Flight Requirements
QE at 0.25 keV								
▪ Bare CCD	0.73	0.8	N/A (FI)	0.8	0.9	0.95	0.95	0.95
▪ CCD+OBF	0.15	0.25	N/A (FI)	0.3	0.8	0.86	0.86	0.86
Device Yield								
▪ FI	0.1	0.8	0.8 est	0.8	0.8	0.8		
▪ BI	0.02	0.25	N/A (FI)	0.25	0.25	0.25		
▪ Net = FI*BI	0.002	0.2		0.2	0.2	0.2		0.2
CCD Frame Rate (Hz)	0.5	2	10	50	50	50	50	50
EDCCD Config		—	FI	FI, BI	FI, BI	BI	BI	BI
Energy Resolution (eV)								
@1.5 keV	130 (S3-BI)	69 (FI)	70(FI) pred	125 (BI)	100 (BI)	100 (BI)	100	100
@0.25 keV	110 (S3-BI)	91 (LTM-BI)	N/A (FI)	125 (BI)	100 (BI)	100 (BI)	100	100
Event Reconstruction	3x3, 5x5	3x3, 5x5	3x3	3x3,5x5	3x3,5x5	3x3,5x5	3x3,5x5	3x3,5x5
Array Format	1024 <sup>2</sup>		512 <sup>2</sup>	1024 <sup>2</sup>	1024 <sup>2</sup>	1024 <sup>2</sup>	1024 <sup>2</sup>	1024 <sup>2</sup>
Focal Plane Complexity	10 chips	48 chips	1 chip	1 chip	1 chip	1 chip	4 chips	13 chips
Radiation Tests			Y	Y		Y	Y	
Environmental Tests							Y	
Milestone Dates	2Q FY97	3Q FY03	3Q FY03	2Q FY03	1Q Fy05	3Q FY05	4Q FY05	
TRL		TRL3		TRL4			TRL 6	
Technology Gate						◆		

## Management, Cost and Schedule

- **Organization Structure**
- **Acquisition Strategies**
- **Schedules**
  - Detailed
  - Critical Paths
- **Cost Estimates**
  - Grass roots, Industry ROMs, PRICE-H
  - Contingency

## Constellation-X Project Organization (Formulation)

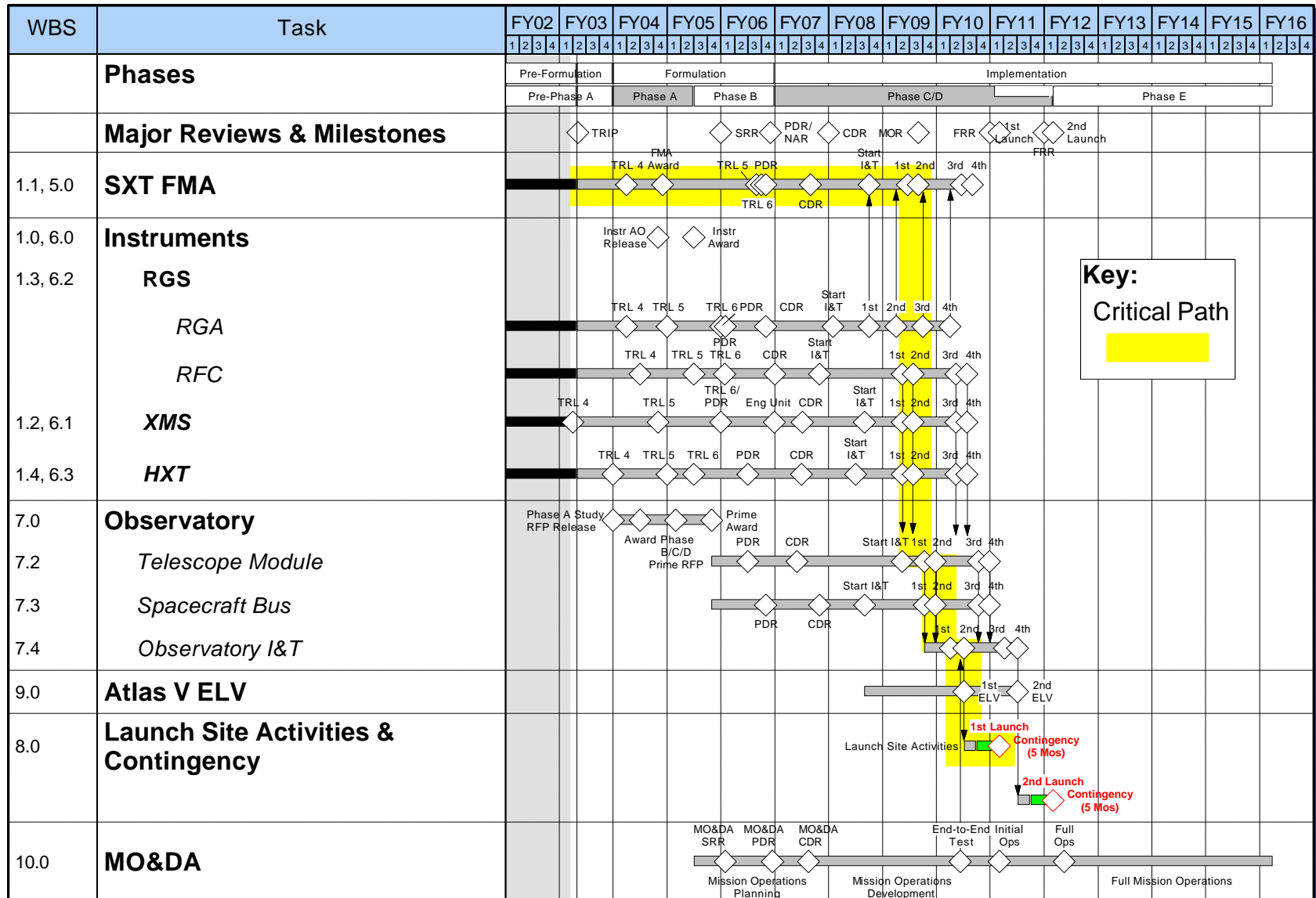




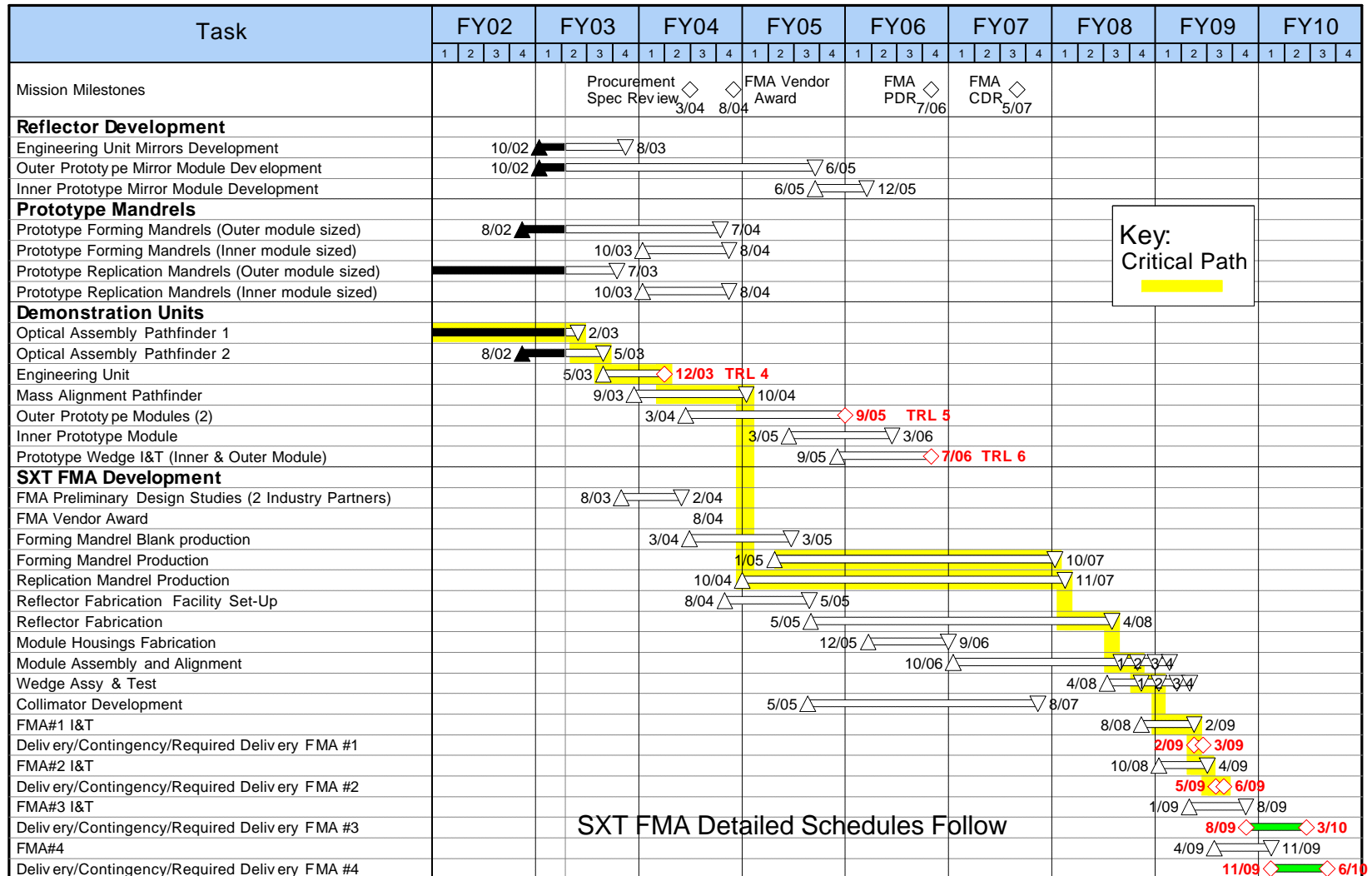
## Acquisition Strategy Summary

System	Solicitation	Contract Award
Instruments (RGS, XMS, HXT)	Announcement of Opportunity	Phase B start
SXT FMA	RFP	Mid-Phase A
Observatory	RFP	Phase B start
Science and Operations Center (CXSOC)	Sole Source to SAO	Phase B start
Ground Stations	RFP for commercial lease	L – 18 months
Launch Vehicle	KSC procurement	L – 30 months

## TRIP Constellation-X Mission Summary Schedule



# SXT Flight Mirror Assembly (FMA) Development Schedule



## **TRIP Review Panel Report Summary—April 22, 2003**

- *GSFC and SAO are a Strong Team, Experienced in X-Ray Astronomy Missions*
- *The Mission has Strong Support from GSFC Management*
- *The Project Benefits from a Rich Heritage While at the same Time Pushing the Envelope in Several Key Technology Areas*
- *The Schedule and Budget Reserves are Low, Especially Early Year Funding*
- *The Review Panel Feels that with Added Budget and Schedule Reserves, Con-X has a High Likelihood of Reaching the Launch Pad Successfully and on Time*

## TRIP Highlights Summary and Conclusions

- **Generation of TRIP report valuable**
  - Overall mission planning taken to next level of detail
  - Report itself is handy reference
- **Independent TRIP Review assessment validates mission concept and plans; identifies areas to further reduce risk**
- **Most of work for TRIP is applicable to stretched out schedule**
- **Project is positioned to enter into Phase A**